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# Complete space-like hypersurfaces in locally symmetric Lorentz spaces 

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#### Abstract

The purpose of this paper is to study complete space-like hypersurfaces with constant mean curvature in a locally symmetric Lorentz space satisfying some curvature conditions. We give an optimal estimate of the squared norm of the second fundamental form of such hypersurfaces. Furthermore, the totally umbilical hypersurfaces are characterized.


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## 1. Introduction

Let $M_{s}^{m}$ be an $m$-dimensional connected semi-Riemannian manifold of index $s(\geq 0)$. It is called a semi-definite space of index $s$. In particular, $M_{1}^{m}$ is called a Lorentz space. A hypersurface $M$ of a Lorentz space is said to be space-like if the induced metric on $M$ from that of the Lorentz space is positive definite. When the Lorentz space $M_{1}^{m}$ is of constant curvature $c$, we call it Lorentz space form, denoted by $M_{1}^{m}(c)$.

It is well known that a maximal space-like entire graph in Minkowski space $\mathbf{R}_{1}^{n+1}$ is a linear hyperplane $[2,7]$. As a generalization of the result above, Chouque-Bruhat et al.

[^0][10] and Ishihara [11] proved that totally geodesic hypersurfaces are the only complete space-like maximal hypersurfaces in $M_{1}^{n+1}(c), c \geq 0$.

On the other hand, it was pointed out by Marsdan and Tipler [13] and Stumbles [19] that space-like hypersurfaces with constant mean curvature in arbitrary spacetime get interested in the relativity theory. Space-like hypersurfaces with constant mean curvature are convenient as initial hypersurfaces for the Cauchy problem in arbitrary spacetime and for studying the propagation of gravitational radiation. Hence, complete space-like hypersurfaces with constant mean curvature in a Lorentz space form $M_{1}^{n+1}(c)$ are extensively investigated by many differential geometers in various view points; for example, Akutagawa [1] and Ramanathan [18] proved complete space-like hypersurfaces with constant mean curvature $H$ in the de Sitter space $S_{1}^{n+1}(c)$ must be totally umbilical if $n^{2} H^{2}<4(n-1) c$ when $n>2$, and $H^{2} \leq 1$ when $n=2$. For further development in these directions, see [3-6,12,14,15] and so on.

As standard models of complete space-like hypersurfaces with constant mean curvature in $M_{1}^{n+1}(c)$, there are four classes of complete hypersurfaces $\mathbf{H}^{k}\left(c_{1}\right) \times S^{n-k}\left(c_{2}\right), \mathbf{R}^{k} \times$ $S^{n-k}\left(c_{2}\right), \mathbf{H}^{k}\left(c_{1}\right) \times \mathbf{R}^{n-k}$ and $\mathbf{H}^{k}\left(c_{1}\right) \times \mathbf{H}^{n-k}\left(c_{2}\right)$, where $k=0,1, \ldots, n$, according to $c>0,=0$ or $<0$. In particular, $\mathbf{H}^{1}\left(c_{1}\right) \times S^{n-1}\left(c_{2}\right)$ is called a hyperbolic cylinder and $\mathbf{H}^{n-1}\left(c_{1}\right) \times S^{1}\left(c_{2}\right)$ is called a spherical cylinder.

It is important and natural to study complete space-like hypersurfaces with constant mean curvature in the more general Lorentz spaces since they have important meaning in the relativity theory. First of all, we shall consider several examples of Lorentz spaces which are not Lorentz space forms.

Example 1. We consider the semi-Riemannian product manifold

$$
H_{1}^{k}\left(-\frac{c_{1}}{n}\right) \times M^{n+1-k}\left(c_{2}\right), c_{1}>0
$$

Its sectional curvature is given by

$$
K^{\prime}\left(u_{1}, u_{b}\right)=-\frac{c_{1}}{n}, \quad K^{\prime}\left(u_{a}, u_{b}\right)=-\frac{c_{1}}{n}, \quad K^{\prime}\left(u_{a}, u_{r}\right)=0, \quad K^{\prime}\left(u_{r}, u_{s}\right)=c_{2}
$$

where $a, b, \ldots=2, \ldots k, r, s, \ldots=k+1, \ldots, n+1$, and $u_{1}$ and $u_{a}, u_{r}$ denotes time-like and space-like vectors respectively.

Example 2. We consider the semi-Riemannian product manifold

$$
R_{1}^{k} \times S^{n+1-k}(1)
$$

Its sectional curvature is given by

$$
K^{\prime}\left(u_{1}, u_{a}\right)=0, \quad K^{\prime}\left(u_{a}, u_{b}\right)=0, \quad K^{\prime}\left(u_{1}, u_{r}\right)=0, \quad K^{\prime}\left(u_{r}, u_{s}\right)=1
$$

where $a, b, \ldots=2, \ldots, k$ and $r, s, \ldots=k+1, \ldots, n+1$. In particular, $R_{1}^{1} \times S^{n}(1)$ is so-called Einstein Static Universe. Of course, it is not Lorentzian space form.

Next, we shall show a general example of Lorentz space which is called Robertson-Walker spacetime.

Example 3. Let $I$ denote an open interval of $R_{1}^{1}$ and $f>0$ a smooth function defined on the interval $I$. For a three-dimensional Riemannian manifold of constant curvature $c$, $c=-1,0,1$, we construct a Lorentz space $M(c, f)$ as the warped product

$$
M(c, f)=I \times_{f} M^{3}(c)
$$

which is called Robertson-Walker spacetime (see [17], pp. 343-345). Thus, by a direct computation, we have that its sectional curvature is given by

$$
K^{\prime}(v, w)=\left(\frac{f^{\prime}}{f}\right)^{2}+\frac{c}{f^{2}}
$$

for any space-like vectors $v$ and $w$ and

$$
K^{\prime}\left(e_{0}, v\right)=\frac{f^{\prime \prime}}{f}
$$

for any time-like vector $e_{0}$ and any unit space-like vector $v$.
In this paper, we shall consider $n+1$-dimensional Lorentz spaces $M^{\prime}$ of index 1. Let $\nabla^{\prime}$, $K^{\prime}$ and $R^{\prime}$ denote the semi-Riemannian connection, sectional curvature and the curvature tensor on $M^{\prime}$, respectively. For constants $c_{1}, c_{2}$ and $c_{3}$, we consider Lorentz spaces which satisfy the following:
(1) for any space-like vector $u$ and any time-like vector $v$

$$
K^{\prime}(u, v)=-\frac{c_{1}}{n}
$$

(2) for any space-like vectors $u$ and $v$

$$
\begin{align*}
& K^{\prime}(u, v) \geq c_{2} \\
& \left|\nabla^{\prime} R^{\prime}\right| \leq \frac{c_{3}}{n} \tag{3}
\end{align*}
$$

When $M^{\prime}$ satisfies conditions (1) and (2), we shall say that $M^{\prime}$ satisfies condition (*). When $M^{\prime}$ satisfies conditions (1)-(3), we shall say that $M^{\prime}$ satisfies condition (**).

Remark 1. It can be easily seen that if the Lorentz space $M^{\prime}$ is locally symmetric, then the condition (3) holds.

Remark 2. The Lorentz space form $M_{1}^{n+1}(c)$ satisfies the conditions ( $*$ ) and ( $* *$ ), where $-\left(c_{1} / n\right)=c_{2}=c$.

Remark 3. The Examples 1 and 2 satisfy the conditions (*) and (**) and the Example 3 also satisfies the conditions $(*)$ and $(* *)$ if we choose an appropriate function $f$ (see [17], pp. 343-345).

In [9,20], authors investigated complete space-like hypersurfaces $M$ in a Lorentz space satisfying condition $(* *)$. They estimated the squared norm of the second fundamental form of $M$ under some conditions. In this paper, we shall prove the following theorem.

Theorem 1. Let $M$ be a complete space-like hypersurface with constant mean curvature $H$ in an $n+1$-dimensional locally symmetric Lorentz space $M^{\prime}$ satisfying the condition (*):
(1) If $n^{2} H^{2}<4(n-1) c$, where $c=2 c_{2}+\left(c_{1} / n\right)$, then $c>0, S \equiv n H^{2}$ and $M$ is totally umbilical, where $S$ denotes the squared norm of the second fundamental form of $M$.
(2) If $n^{2} H^{2}=4(n-1) c$, then $c \geq 0$ and either $S \equiv n H^{2}$ and $M$ is totally umbilical, or $\sup S=n c$.
(3) If $n^{2} H^{2}>4(n-1) c$ and $c<0$, then either $S \equiv n H^{2}$ and $M$ is totally umbilical, or

$$
n H^{2}<\sup S \leq S_{\max }
$$

where $S_{\max }=n / 2(n-1)\left[n^{2} H^{2}-2(n-1) c+(n-2)|H|\left\{n^{2} H^{2}-4(n-1) c\right\}^{1 / 2}\right]$.
(4) If $n^{2} H^{2}>4(n-1) c$ and $c \geq 0$, then either $S \equiv n H^{2}$ and $M$ is totally umbilical, or

$$
S_{\max } \geq \sup S \begin{cases}>n H^{2}, & \text { if } H^{2} \geq c \\ \geq S_{\min }, & \text { if } H^{2}<c\end{cases}
$$

where $S_{\min }=n / 2(n-1)\left[n^{2} H^{2}-2(n-1) c-(n-2)|H|\left\{n^{2} H^{2}-4(n-1) c\right\}^{1 / 2}\right]$.

$$
\begin{equation*}
S \equiv \frac{n}{2(n-1)}\left[n^{2} H^{2}-2(n-1) c+(n-2)|H|\left\{n^{2} H^{2}-4(n-1) c\right\}^{1 / 2}\right] \tag{5}
\end{equation*}
$$

if and only if $M$ is an isoparametric hypersurface with two distinct principal curvatures one of which is simple.

Theorem 2. Let $M$ be an $n(n>2)$ dimensional complete space-like hypersurface with constant mean curvature $H$ in an $n+1$-dimensional locally symmetric Lorentz space $M^{\prime}$ satisfying the condition $(*)$. If the sectional curvature of $M$ is not less than $-\left(c_{2}+\left(c_{1} / n\right)\right)$, then, $c \geq 0$, where $c=2 c_{2}+\left(c_{1} / n\right)$. Furthermore, if $H^{2} \geq c$ and

$$
S<\frac{n}{2(n-1)}\left[n^{2} H^{2}-2(n-1) c+(n-2)|H|\left\{n^{2} H^{2}-4(n-1) c\right\}^{1 / 2}\right]
$$

hold, then $M$ is totally umbilical.
Remark 4. We should notice that when $M^{\prime}$ is a Lorentz space form $M_{1}^{n+1}(c)$, a part of the similar results in Theorems 1 and 2 was obtained by Cheng and Nakagawa [6] and Ki et al. [12].

Remark 5. Euclidean space $R^{n}$ which is defined by $x_{1}=x_{n+2}+t$ is a totally umbilical space-like hypersurface of $S_{1}^{n+1}(c)$ in $R_{1}^{n+2}$, where $\left\{x_{1}, \ldots, x_{n+2}\right\}$ is the natural coordinate system in $R_{1}^{n+2}$. The mean curvature $H$ satisfies $H^{2}=c$.

Remark 6. We consider a family of space-like hypersurfaces $H^{k}\left(c_{1}\right) \times S^{n-k}\left(c_{2}\right)$ of $S_{1}^{n+1}(c)$ which is defined by

$$
\begin{aligned}
& H^{k}\left(c_{1}\right) \times S^{n-k}\left(c_{2}\right) \\
& \quad=\left\{(x, y) \in S_{1}^{n+1}(c) \subset R_{1}^{n+2}=R_{1}^{k+1} \times R^{n-k+1}:|x|^{2}=-\frac{1}{c_{1}},|y|^{2}=\frac{1}{c_{2}}\right\}
\end{aligned}
$$

where $c_{1}<0, c_{2}>0$ and $1 / c_{1}+1 / c_{2}=1 / c$. When $k>1$, it is not of non-negative curvature. The number of distinct principal curvatures of such a hypersurface is exactly two. A principal curvature is equal to $\left(c-c_{1}\right)^{1 / 2}$ with the multiplicity $k$ and the other is equal to $\left(c-c_{2}\right)^{1 / 2}$ with multiplicity $n-k$. We can prove that

$$
S<\frac{1}{2(n-1)}\left[n\left\{n^{2} H^{2}-2(n-1) c\right\}+(n-2) n|H|\left\{n^{2} H^{2}-4(n-1) c\right\}^{1 / 2}\right],
$$

and $H^{2} \geq c$ if we choose appropriate values $c_{1}$ and $c_{2}$. Hence, the assumption on sectional curvature in Theorem 2 is essential.

Remark 7. Hyperbolic cylinder $H^{1}\left(c_{1}\right) \times S^{n-1}\left(c_{2}\right)$ has non-negative curvature and satisfies $H^{2} \geq c$ if we choose appropriate values $c_{1}$ and $c_{2}$. The squared norm $S$ of the second fundamental form $\alpha$ satisfies $S=1 / 2(n-1)\left[n\left\{n^{2} H^{2}-2(n-1) c\right\}+(n-2) n|H|\left\{n^{2} H^{2}-\right.\right.$ $4(n-1) c\}^{1 / 2}$. Hence, the estimate of $S$ in Theorem 2 is best possible.

## 2. Preliminaries

First of all, we review basic formulas on space-like hypersurfaces in a Lorentz space. Let ( $M^{\prime}, g^{\prime}$ ) be an $(n+1)$-dimensional Lorentz space, i.e., an indefinite Riemannian manifold of index 1. Throughout this paper, manifolds are always assumed to be connected and geometric objects are assumed to be of class $C^{\infty}$. For any point $x$ in $M^{\prime}$ we choose a local field of orthonormal frames $\left\{e_{A}\right\}=\left\{e_{0}, e_{1}, \ldots, e_{n}\right\}$ on a neighborhood of $x$. Here and in the sequel, the following convention on the range of indices will be used throughout this paper, unless otherwise stated:

$$
A, B, \ldots=0,1, \ldots, n, \quad i, j, \ldots=1, \ldots, n
$$

Let $\left\{\omega_{A}\right\}=\left\{\omega_{0}, \omega_{1}, \ldots, \omega_{n}\right\}$ denote the dual frame fields of $\left\{e_{A}\right\}$ on $M^{\prime}$. Then metric tensor $g^{\prime}$ of $M^{\prime}$ satisfies $g^{\prime}\left(e_{A}, e_{B}\right)=\epsilon_{A} \delta_{A B}$, where $\epsilon_{0}=-1$ and $\epsilon_{j}=1$. The canonical forms $\omega_{A}$ and the connection forms $\omega_{A B}$ of $M^{\prime}$ satisfy the structure equations:

$$
\begin{align*}
& \mathrm{d} \omega_{A}+\sum \epsilon_{B} \omega_{A B} \wedge \omega_{B}=0, \quad \omega_{A B}+\omega_{B A}=0,  \tag{2.1}\\
& \mathrm{~d} \omega_{A B}+\sum \epsilon_{C} \omega_{A C} \wedge \omega_{C B}=\Omega_{A B}^{\prime},  \tag{2.2}\\
& \Omega_{A B}^{\prime}=-\frac{1}{2} \sum \epsilon_{C} \epsilon_{D} R_{A B C D}^{\prime} \omega_{C} \wedge \omega_{D}, \tag{2.3}
\end{align*}
$$

where $\Omega^{\prime}=\Omega_{A B}^{\prime}$ (resp. $R_{A B C D}^{\prime}$ ) denotes the Riemannian curvature form (resp. the components of the Riemannian curvature tensor $R^{\prime}$ ) of $M^{\prime}$. The components $R_{C D}^{\prime}$ of the Ricci tensor and the scalar curvature $r^{\prime}$ are given by

$$
\begin{align*}
& R_{C D}^{\prime}=\sum_{B} \epsilon_{B} R_{B C D B}^{\prime}  \tag{2.4}\\
& r^{\prime}=\sum_{A} \epsilon_{A} R_{A A}^{\prime} \tag{2.5}
\end{align*}
$$

respectively. The components $R_{A B C D ; E}^{\prime}$ of the covariant derivative of the Riemannian curvature tensor $R^{\prime}$ are defined by

$$
\begin{align*}
& \sum_{E} \epsilon_{E} R_{A B C D ; E}^{\prime} \omega_{E} \\
& \quad=\mathrm{d} R_{A B C D}^{\prime}-\sum_{E} \epsilon_{E}\left(R_{E B C D}^{\prime} \omega_{E A}+R_{A E C D}^{\prime} \omega_{E B}+R_{A B E D}^{\prime} \omega_{E C}+R_{A B C E}^{\prime} \omega_{E D}\right) \tag{2.6}
\end{align*}
$$

A plane section $P^{\prime}$ of the tangent space $T_{x} M^{\prime}$ of $M^{\prime}$ at any point $x$ is said to be non-degenerate, provided that $g_{x} \mid P^{\prime}$ is non-degenerate. It is easily seen that $P^{\prime}$ is non-degenerate if and only if it has a basis $\{u, v\}$ such that $g(u, u) g(v, v)-g(u, v)^{2} \neq 0$. The sectional curvature of the non-degenerate plane section $P^{\prime}$ spanned by $u$ and $v$ is denoted by $K^{\prime}\left(P^{\prime}\right)=K^{\prime}(u, v)$. The semi-definite Riemannian manifold $M^{\prime}$ is said to be of constant curvature if its sectional curvature $K^{\prime}\left(P^{\prime}\right)$ is constant for all $P^{\prime}$ and for all points of $M^{\prime} . M^{\prime}$ is called a semi-definite space form if it is of constant curvature. An $m$-dimensional semi-definite space form of constant curvature $c$ and of index $s$ is denoted by $M_{s}^{m}(c)$. The standard models of semi-definite space forms are the following three kinds: the semi-definite Euclidean space $R_{s}^{m}$, the semi-definite spherical space $S_{s}^{m}(c)$ or the semi-definite hyperbolic space $H_{s}^{m}(c)$, according to $c=0,>0$ or $<0$. The Riemannian curvature tensor $R_{A B C D}$ of the semi-definite space form $M_{s}^{m}(c)$ is given by

$$
\begin{equation*}
R_{A B C D}=c \epsilon_{A} \epsilon_{B}\left(\delta_{A D} \delta_{B C}-\delta_{A C} \delta_{B D}\right) \tag{2.7}
\end{equation*}
$$

Now, let ( $M^{\prime}, g^{\prime}$ ) be an $(n+1)$-dimensional Lorentz space and let $M$ be an $n$-dimensional space-like hypersurface of $M^{\prime}$. We choose a local field of orthonormal frames $\left\{e_{A}\right\}=$ $\left\{e_{0}, e_{1} \ldots, e_{n}\right\}$ in such a way that restricted to $M, e_{1}, \ldots, e_{n}$ are tangent to $M$ and the other is normal to $M$. Namely, $e_{1}, \ldots, e_{n}$ are space-like vectors and the other $e_{0}$ is a time-like vector. Let $\left\{\omega_{A}\right\}$ be its dual frame field. Then the indefinite Riemannian metric tensor $g^{\prime}$ of $M^{\prime}$ is given by $g^{\prime}=\sum_{A} \epsilon_{A} \omega_{A} \otimes \omega_{A}$. The connection forms on $M^{\prime}$ are denoted by $\omega_{A B}$, where $\epsilon_{0}=-1$ and $\epsilon_{j}=1$.

Restricting these forms to the space-like hypersurface $M$ in $M^{\prime}$, we have

$$
\begin{equation*}
\omega_{0}=0 \tag{2.8}
\end{equation*}
$$

and the induced metric $g$ of $M$ is given by $g=\sum \omega_{j} \otimes \omega_{j}$. From (2.1) and (2.8) and the Cartan lemma, we have

$$
\begin{equation*}
\omega_{0 i}=\sum_{j} h_{i j} \omega_{j}, h_{i j}=h_{j i} \tag{2.9}
\end{equation*}
$$

The quadratic form $\alpha=-\sum_{i, j} h_{i j} \omega_{i} \otimes \omega_{j} \otimes e_{0}$ with values in the normal bundle and $H=$ $1 / n \sum_{j=1}^{n} h_{j j}$ are called second fundamental form and mean curvature of the hypersurface $M$, respectively. When principal curvatures of $M$ are constant, $M$ is called isoparametric. The connection forms $\left\{\omega_{i j}\right\}$ of $M$ are characterized by the structure equation of $M$ :

$$
\begin{align*}
\mathrm{d} \omega_{i}+\sum_{j} \omega_{i j} \wedge \omega_{j} & =0, \quad \omega_{i j}+\omega_{j i}
\end{aligned}=0, \quad \begin{aligned}
& \mathrm{d} \omega_{i j}+\sum_{k} \omega_{i k} \wedge \omega_{k j}=\Omega_{i j}, \quad \frac{1}{2} \sum_{k, l} R_{i j k l} \omega_{k} \wedge \omega_{l} \tag{2.10}
\end{align*}
$$

From (2.3) and (2.11), we have Gauss equation

$$
\begin{equation*}
R_{i j k l}=R_{i j k l}^{\prime}-\left(h_{i l} h_{j k}-h_{i k} h_{j l}\right) \tag{2.12}
\end{equation*}
$$

Components $R_{i j}$ of Ricci tensor and scalar curvature $r$ of $M$ are given by

$$
\begin{align*}
& R_{i j}=\sum_{k=1}^{n} R_{k i j k}^{\prime}-n H h_{i j}+\sum_{k=1}^{n} h_{i k} h_{k j}  \tag{2.13}\\
& r=\sum_{j, k=1}^{n} R_{k j j k}^{\prime}-n^{2} H^{2}+S \tag{2.14}
\end{align*}
$$

where $S=\sum_{i, j=1}^{n} h_{i j}^{2}$ denotes the squared norm of the second fundamental form of $M$.
By taking exterior differentiation of (2.9) and defining $h_{i j k}$ by

$$
\begin{equation*}
\sum_{k} h_{i j k} \omega_{k}=\mathrm{d} h_{i j}-\sum_{k}\left(h_{k j} \omega_{k i}+h_{i k} \omega_{k j}\right) \tag{2.15}
\end{equation*}
$$

we have Codazzi equation

$$
\begin{equation*}
h_{i j k}-h_{i k j}=R_{0 i j k}^{\prime} \tag{2.16}
\end{equation*}
$$

Similarly, defining $h_{i j k l}$ by

$$
\begin{equation*}
\sum_{l} h_{i j k l} \omega_{l}=\mathrm{d} h_{i j k}-\sum_{l}\left(h_{l j k} \omega_{l i}+h_{i l k} \omega_{l j}+h_{i j l} \omega_{l k}\right) \tag{2.17}
\end{equation*}
$$

and differentiating (2.15) exteriorly, we have

$$
\begin{aligned}
\sum_{k} & \left\{\mathrm{~d} h_{i j k} \wedge \omega_{k}+h_{i j k}\left(-\sum_{l} \omega_{k l} \wedge \omega_{l}\right)\right\} \\
= & -\sum_{k}\left[\mathrm{~d} h_{k j} \wedge \omega_{k i}+h_{j k}\left\{-\sum_{l} \omega_{k l} \wedge \omega_{l i}-\frac{1}{2} \sum_{l, m} R_{k i l m} \omega_{l} \wedge \omega_{m}\right\}\right. \\
& \left.+\mathrm{d} h_{i k} \wedge \omega_{k j}+h_{i k}\left\{-\sum_{l} \omega_{k l} \wedge \omega_{l j}-\frac{1}{2} \sum_{l, m} R_{k j l m} \omega_{l} \wedge \omega_{m}\right\}\right]
\end{aligned}
$$

Hence, we obtain Ricci formula for the second fundamental form of $M$ :

$$
\begin{equation*}
h_{i j k l}-h_{i j l k}=-\sum_{r}\left(h_{i r} R_{r j k l}+h_{j r} R_{r i k l}\right) \tag{2.18}
\end{equation*}
$$

Now let us denote, by $R_{A B C D ; E}^{\prime}$, covariant derivative of $R_{A B C D}^{\prime}$. Then, restricting on $M$, $R_{0 i j k ; l}^{\prime}$ is given by

$$
\begin{equation*}
R_{0 i j k ; l}^{\prime}=R_{0 i j k l}^{\prime}+R_{0 i 0 k}^{\prime} h_{j l}+R_{0 i j 0}^{\prime} h_{k l}+\sum_{m} R_{m i j k}^{\prime} h_{m l}, \tag{2.19}
\end{equation*}
$$

where $R_{0 i j k l}^{\prime}$ denote the covariant derivative of $R_{0 i j k}^{\prime}$ as a tensor on $M$ so that

$$
\sum_{l} R_{0 i j k l}^{\prime} \omega_{l}=\mathrm{d} R_{0 i j k}^{\prime}-\sum_{l} R_{0 l j k}^{\prime} \omega_{l i}-\sum_{l} R_{0 i l k}^{\prime} \omega_{l j}-\sum_{l} R_{0 i j l}^{\prime} \omega_{l k} .
$$

Next, we compute the Laplacian $\Delta h_{i j}$ defined by

$$
\begin{equation*}
\Delta h_{i j}=\sum_{k} h_{i j k k} \tag{2.20}
\end{equation*}
$$

From (2.16) and (2.18) it follows that

$$
\begin{aligned}
\Delta h_{i j} & =\sum_{k} h_{i k j k}+\sum_{k} R_{0 i j k k}^{\prime}=\sum_{k} h_{k i j k}+\sum_{k} R_{0 i j k k}^{\prime} \\
& =\sum_{k}\left\{h_{k i k j}-\sum_{l}\left(h_{k l} R_{l i j k}+h_{i l} R_{l k j k}\right)+R_{0 i j k k}^{\prime}\right\} .
\end{aligned}
$$

From $h_{k i k j}=h_{k k i j}+R_{0 k i k j}^{\prime}$, we obtain

$$
\Delta h_{i j}=\sum_{k} h_{k k i j}+\sum_{k}\left(R_{0 i j k k}^{\prime}+R_{0 k i k j}^{\prime}\right)-\sum_{k, l}\left(h_{k l} R_{l i j k}+h_{i l} R_{l j j k}\right) .
$$

By (2.12) and (2.19) and the above equation, we obtain

$$
\begin{align*}
\Delta h_{i j}= & \sum_{k} h_{k k i j}+\sum_{k}\left(R_{0 i j k ; k}^{\prime}+R_{0 k i k ; j}^{\prime}\right)-\sum_{k}\left(h_{j k} R_{0 i 0 k}^{\prime}+h_{k k} R_{0 i j 0}^{\prime}+\sum_{l} h_{k l} R_{l i j k}^{\prime}\right) \\
& -\sum_{k}\left(h_{i j} R_{0 k 0 k}^{\prime}+h_{k j} R_{0 k i 0}^{\prime}+\sum_{l} h_{j l} R_{l k i k}^{\prime}\right)-\sum_{k, l}\left(R_{l k j k}^{\prime}-h_{l k} h_{j k}+h_{k k} h_{j l}\right) h_{i l} \\
& -\sum_{k, l}\left(R_{l i j k}^{\prime}-h_{k l} h_{i j}+h_{l j} h_{i k}\right) h_{k l} \\
= & \sum_{k} h_{k k i j}+\sum_{k}\left(R_{0 i j k ; k}^{\prime}+R_{0 k i k ; j}^{\prime}\right)-\sum_{k}\left(h_{k k} R_{0 i j 0}^{\prime}+h_{i j} R_{0 k 0 k}^{\prime}\right) \\
& -\sum_{k, l}\left(2 h_{k l} R_{l i j k}^{\prime}+h_{j l} R_{l k i k}^{\prime}+h_{i l} R_{l k j k}^{\prime}\right)-n H \sum_{l} h_{i l} h_{l j}+S h_{i j} . \tag{2.21}
\end{align*}
$$

The following Generalized Maximum Principle of Omori [16] and Yau [21] will play an important role in the proof of our Theorems (cf. [8,20]).

Generalized Maximum Principle. Let $M$ be a complete Riemannian manifold whose Ricci curvature is bounded from below on $M$. Let $F$ be a $C^{2}$-function bounded from above on $M$, then, for any $\epsilon>0$, there exists a point $p \in M$ such that

$$
\begin{equation*}
|\nabla F(p)|<\epsilon, \quad \Delta F(p)<\epsilon \quad \text { and } \quad \sup F-\epsilon<F(p) \tag{2.22}
\end{equation*}
$$

## 3. Locally formulas

In this section, we assume that $M^{\prime}$ is an $(n+1)$-dimensional Lorentz space satisfying condition ( $* *$ ) and $M$ is a space-like hypersurface with constant mean curvature in $M^{\prime}$. First of all, we calculate the Laplacian of the squared norm $S$ of the second fundamental form $\alpha$ of $M$.

$$
\begin{equation*}
\frac{1}{2} \Delta S=\frac{1}{2} \Delta\left(\sum_{i, j} h_{i j}^{2}\right)=\sum_{i, j, k}\left(h_{i j k} h_{i j}\right)_{k}=\sum_{i, j, k} h_{i j k}^{2}+\sum_{i, j, k} h_{i j k k} h_{i j} \tag{3.1}
\end{equation*}
$$

By (2.21), we have

$$
\begin{align*}
\frac{1}{2} \Delta S= & \sum_{i, j, k} h_{i j k}^{2}+\sum_{i, j}\left[\sum_{k}\left(R_{0 i j 0 ; j}^{\prime}+R_{0 i j k ; k}^{\prime}\right)-\sum_{k}\left(h_{k k} R_{0 i j 0}^{\prime}+h_{i j} R_{0 k 0 k}^{\prime}\right)\right. \\
& \left.-\sum_{k, l}\left(2 h_{k l} R_{l i j k}^{\prime}+h_{l j} R_{l k i k}^{\prime}+h_{l i} R_{k j k l}^{\prime}\right)-n H \sum_{l} h_{i l} h_{l j}+S h_{i j}\right] h_{i j} \tag{3.2}
\end{align*}
$$

Thus, we have

$$
\begin{align*}
\frac{1}{2} \Delta S= & \sum_{i, j, k} h_{i j k}^{2}+\sum_{i, j, k} h_{i j}\left(R_{0 k i k ; j}^{\prime}+R_{0 i j k ; k}^{\prime}\right)-\left(\sum_{i, j} n H h_{i j} R_{0 j i 0}^{\prime}+S \sum_{k} R_{0 k 0 k}^{\prime}\right) \\
& -\sum_{i, j, k, l} 2\left(h_{i j} h_{k l} R_{l i j k}^{\prime}+h_{l i} h_{i j} R_{l k j k}^{\prime}\right)-n H h_{3}+S^{2} \tag{3.3}
\end{align*}
$$

where $h_{3}=\sum_{j=1}^{n} \lambda_{j}^{3}$ and $\lambda_{j}$ 's are principal curvatures of $M$.
Next, we will choose $\left\{e_{1}, \ldots, e_{n}\right\}$ such that

$$
\begin{equation*}
h_{i j}=\lambda_{i} \delta_{i j} \tag{3.4}
\end{equation*}
$$

By definition, we see

$$
S=\sum_{i} \lambda_{i}^{2}
$$

Putting $\mu_{j}=\lambda_{j}-H$, we have

$$
\begin{equation*}
\sum_{j} \mu_{j}=0, \quad \sum_{j} \mu_{j}^{2}=S-n H^{2} \tag{3.5}
\end{equation*}
$$

Therefore, for any $j$

$$
\begin{equation*}
\left(\lambda_{j}-H\right)^{2} \leq \frac{n-1}{n}\left(S-n H^{2}\right) \tag{3.6}
\end{equation*}
$$

## Since

$$
\begin{aligned}
& \sum_{i, j, k}\left(R_{0 k i k ; j}^{\prime}+R_{0 i j k ; k}^{\prime}\right) h_{i j} \\
& \quad=\sum_{j, k} \lambda_{j}\left(R_{0 k j k ; j}^{\prime}+R_{0 j j k ; k}^{\prime}\right) \\
& \quad \geq-\sum_{k}\left[|H|+\sqrt{\frac{n-1}{n}\left(S-n H^{2}\right)}\right]\left(\sqrt{\sum_{j} R_{0 k j k ; j}^{\prime 2}}+\sqrt{\sum_{j} R_{0 j j k ; k}^{\prime 2}}\right) \\
& \quad \geq-2 \sqrt{n}\left[|H|+\sqrt{\frac{n-1}{n}\left(S-n H^{2}\right)}\right]\left|\nabla R^{\prime}\right|,
\end{aligned}
$$

we obtain

$$
\begin{equation*}
\sum_{i, j, k}\left(R_{0 k i k ; j}^{\prime}+R_{0 i j k ; k}^{\prime}\right) h_{i j} \geq-\frac{2}{\sqrt{n}}\left[|H|+\sqrt{\frac{n-1}{n}\left(S-n H^{2}\right)}\right] c_{3} . \tag{3.7}
\end{equation*}
$$

By (3.4) and condition (*), we have

$$
\begin{aligned}
& -\left(\sum_{i, j} n H h_{i j} R_{0 i j 0}^{\prime}+S \sum_{k} R_{0 k 0 k}^{\prime}\right) \\
& \quad=-\sum_{k} n H \lambda_{k}\left(R_{0 k k 0}^{\prime}-S \sum_{k} R_{0 k k 0}^{\prime}\right)=\sum_{k}\left(S-n H \lambda_{k}\right) R_{0 k k 0}^{\prime}=\sum_{k}\left(S-n H \lambda_{k}\right) \frac{c_{1}}{n} .
\end{aligned}
$$

Hence it follows that

$$
\begin{equation*}
-\left(\sum_{i, j} n H h_{i j} R_{0 i j 0}^{\prime}+S \sum_{k} R_{0 k 0 k}^{\prime}\right)=c_{1}\left(S-n H^{2}\right) \tag{3.8}
\end{equation*}
$$

Since

$$
\begin{aligned}
& -\sum_{i, j, k, l}\left(h_{i j} h_{k l} R_{l i j k}^{\prime}+h_{l i} h_{i j} R_{l k j k}^{\prime}\right) \\
& \quad=-\sum_{j, k}\left(\lambda_{j} \lambda_{k} R_{k j j k}^{\prime}-\lambda_{k}^{2} R_{k j j k}^{\prime}\right)=-\sum_{j, k}\left(\lambda_{j} \lambda_{k}-\lambda_{k}^{2}\right) R_{k j j k}^{\prime} \\
& \quad=\frac{1}{2} \sum_{j, k}\left(\lambda_{j}-\lambda_{k}\right)^{2} R_{k j j k}^{\prime} \geq \frac{c_{2}}{2} \sum_{j, k}\left(\lambda_{j}-\lambda_{k}\right)^{2}
\end{aligned}
$$

we obtain

$$
\begin{equation*}
-\sum_{i, j, k, l}\left(h_{i j} h_{k l} R_{l i j k}^{\prime}+h_{l i} h_{i j} R_{l k j k}^{\prime}\right) \geq \frac{c_{2}}{2} \sum_{j, k}\left(\lambda_{j}-\lambda_{k}\right)^{2}=c_{2}\left(n S-n^{2} H^{2}\right) \tag{3.9}
\end{equation*}
$$

Thus, substituting (3.7)-(3.9) into (3.3), we can prove the following lemma.

Lemma 1. Let $M^{\prime}$ be an $(n+1)$-dimensional Lorentz space satisfying the condition $(* *)$. If $M$ is a space-like hypersurface with constant mean curvature $H$ in $M^{\prime}$, then we have

$$
\begin{align*}
\frac{1}{2} \Delta S \geq & -\frac{2}{\sqrt{n}}\left[|H|+\sqrt{\frac{n-1}{n}\left(S-n H^{2}\right)}\right] c_{3} \\
& +\left(2 n c_{2}+c_{1}\right)\left(S-n H^{2}\right)+\left(S^{2}-n H h_{3}\right) \tag{3.10}
\end{align*}
$$

In particular, if $M^{\prime}$ is locally symmetric, we have

$$
\begin{equation*}
\frac{1}{2} \Delta S \geq\left(2 n c_{2}+c_{1}\right)\left(S-n H^{2}\right)+\left(S^{2}-n H h_{3}\right) \tag{3.11}
\end{equation*}
$$

## 4. Proofs of Theorems

This section presents proofs of our theorems.
Proof of Theorem 1. Since $M^{\prime}$ is an $(n+1)$-dimensional locally symmetric Lorentz space satisfying the condition $(*)$, that is, for constants $c_{1}$ and $c_{2}$, we have

$$
\begin{equation*}
K^{\prime}(u, v)=-\frac{c_{1}}{n} \tag{4.1}
\end{equation*}
$$

for any space-like vector $u$ and any time-like vector $v$ and

$$
\begin{equation*}
K^{\prime}\left(u_{1}, u_{2}\right) \geq c_{2} \tag{4.2}
\end{equation*}
$$

for any space-like vectors $u_{1}$ and $u_{2}$. Hence, we have, from (3.11)

$$
\begin{equation*}
\frac{1}{2} \Delta S \geq\left(2 n c_{2}+c_{1}\right)\left(S-n H^{2}\right)+\left(S^{2}-n H h_{3}\right) \tag{4.3}
\end{equation*}
$$

Let $B=\sum_{i} \mu_{i}^{2}$ and $B_{3}=\sum_{i} \mu_{i}^{3}$. We have

$$
\begin{equation*}
B=S-n H^{2}, \quad B_{3}=h_{3}-3 H B-n H^{3} \tag{4.4}
\end{equation*}
$$

Hence, we have

$$
\begin{equation*}
\frac{1}{2} \Delta B \geq\left(2 n c_{2}+c_{1}\right) B+\left(B+n H^{2}\right)^{2}-n H\left(B_{3}+3 H B+n H^{3}\right) \tag{4.5}
\end{equation*}
$$

because $H$ is constant.
Let $a_{1}, \ldots, a_{n}$ be real numbers satisfying $\sum_{i} a_{i}=0$ and $\sum_{i} a_{i}^{2}=B$, then we can prove

$$
\begin{equation*}
\left|\sum_{i} a_{i}^{3}\right| \leq \frac{n-2}{\sqrt{n(n-1)}} B^{3 / 2} \tag{4.6}
\end{equation*}
$$

and the equality holds if and only if at least $n-1$ of the $a_{i}$ 's are equal.
Therefore, we infer

$$
\begin{equation*}
\frac{1}{2} \Delta B \geq B\left\{B-\frac{n-2}{\sqrt{n(n-1)}} n|H| B^{1 / 2}+n c-n H^{2}\right\} \tag{4.7}
\end{equation*}
$$

where $c=2 c_{2}+\left(c_{1} / n\right)$. Since $M^{\prime}$ satisfies the condition $(*)$ and $M$ has constant mean curvature, from (2.13), we know that the Ricci curvature of $M$ is bounded from below. Since we do not know whether $S$ is bounded yet, we consider a function $F$ defined by, for any positive constant $a, F=-1 /(\sqrt{B+a})$. We know that $F$ is bounded because of $B \geq 0$. According to the Generalized Maximum Principle of Omori [16] and Yau [21] in Section 2, for any $\epsilon_{m}>0$, there exists a point $p_{m} \in M$ such that

$$
\begin{equation*}
\Delta F\left(p_{m}\right)<\epsilon_{m}, \quad|\nabla F|\left(p_{m}\right)<\epsilon_{m}, \quad \sup F-\epsilon_{m}<F\left(p_{m}\right) \tag{4.8}
\end{equation*}
$$

Since

$$
\nabla_{i} F=\frac{1}{2} \frac{\nabla_{i} B}{(B+a)^{3 / 2}}, \quad \Delta F=\frac{1}{2} \frac{\Delta B}{(B+a)^{3 / 2}}-\frac{3}{4} \frac{|\nabla B|^{2}}{(B+a)^{5 / 2}},
$$

we have

$$
|\nabla F|=\frac{1}{2}|F|^{3}|\nabla B|,
$$

and

$$
\frac{1}{2}|F|^{4} \Delta B=|F| \Delta F+3|\nabla F|^{2}
$$

From (4.8), we infer

$$
\frac{1}{2}\left|F\left(p_{m}\right)\right|^{4} \Delta B\left(p_{m}\right)<\left|F\left(p_{m}\right)\right| \epsilon_{m}+3 \epsilon_{m}^{2}
$$

For any positive constant $0<\epsilon<1$, letting $\epsilon_{m} \rightarrow 0$, we know that there exists a positive integer $m_{0}$ such that when $m>m_{0},\left|F\left(p_{m}\right)\right| \epsilon_{m}+3 \epsilon_{m}^{2}<\epsilon$ because $F$ is a bounded function. According to (4.7) and the above inequalities, we infer

$$
(1-\epsilon) B^{2}\left(p_{m}\right)-\frac{n-2}{\sqrt{n(n-1)}} n|H| B^{3 / 2}\left(p_{m}\right)+\left(n c-n H^{2}-2 a \epsilon\right) B\left(p_{m}\right)-a^{2} \epsilon<0
$$

Thus, we know that $\left\{B\left(p_{m}\right)\right\}$ is a bounded sequence. Since $\lim _{m \rightarrow \infty} F\left(p_{m}\right)=\sup F=$ $-\inf (1 / \sqrt{B+a})=-(1 / \sqrt{\sup B+a})$, we have $\lim _{m \rightarrow \infty} B\left(p_{m}\right)=\sup B$ from the definition of $F$. Hence, $B$ is bounded. Since $a$ and $\epsilon$ are any positive constants, we infer

$$
\begin{equation*}
\sup B\left\{\sup B-\frac{n-2}{\sqrt{n(n-1)}} n|H| \sup B^{1 / 2}+n c-n H^{2}\right\} \leq 0 \tag{4.9}
\end{equation*}
$$

If $n^{2} H^{2}<4(n-1) c$ holds, then we have $c>0$ and

$$
\sup B-\frac{n-2}{\sqrt{n(n-1)}} n|H| \sup B^{1 / 2}+n c-n H^{2}>0
$$

Hence, we obtain sup $B=0$, that is, $B \equiv 0$. Thus, we infer that $S \equiv n H^{2}$ and $M$ is totally umbilical.

If $n^{2} H^{2}=4(n-1) c$ holds, then we have $c \geq 0$ and

$$
\begin{aligned}
& \sup B-\frac{n-2}{\sqrt{n(n-1)}} n|H| \sup B^{1 / 2}+n c-n H^{2} \\
& \quad=\left(\sup B^{1 / 2}-\frac{n-2}{2 \sqrt{n(n-1)}} n|H|\right)^{2} \geq 0
\end{aligned}
$$

Hence, from (4.9), we have $\sup B=0$, that is, $B \equiv 0$ if $\sup B^{1 / 2} \neq(n-2) /(2 \sqrt{n(n-1)}) n$ $|H|$. Thus, we have that either $B \equiv 0$, that is, $S \equiv n H^{2}$ and $M$ is totally umbilical, or $\sup B^{1 / 2}=(n-2) /(2 \sqrt{n(n-1)}) n|H|$, namely, $\sup S=n c$. Thus, we complete the proof of (2) in Theorem 1.

If $n^{2} H^{2}>4(n-1) c$ and $c<0$ hold, we know

$$
\begin{gathered}
\sup B-\frac{n-2}{\sqrt{n(n-1)}} n|H| \sup B^{1 / 2}+n c-n H^{2} \\
=\left(\sup B^{1 / 2}-B_{\min }^{1 / 2}\right)\left(\sup B^{1 / 2}-B_{\max }^{1 / 2}\right),
\end{gathered}
$$

where $B_{\text {min }}^{1 / 2}=\left((n-2) n|H|-n \sqrt{n^{2} H^{2}-4(n-1) c}\right) / 2 \sqrt{n(n-1)}$ and $B_{\text {max }}^{1 / 2}=$ $\left((n-2) n|H|+n \sqrt{n^{2} H^{2}-4(n-1) c}\right) / 2 \sqrt{n(n-1)}$. Since $B_{\min }^{1 / 2}=((n-2) n|H|$ $\left.-n \sqrt{n^{2} H^{2}-4(n-1) c}\right) / 2 \sqrt{n(n-1)}<0$ holds when $c<0$, from (4.9), we infer either $\sup B=0$, in this case, $M$ is totally umbilical, or $0<\sup B^{1 / 2} \leq((n-2) n|H|+$ $\left.n \sqrt{n^{2} H^{2}-4(n-1) c}\right) / 2 \sqrt{n(n-1)}$. Hence, we know that the assertion (3) in Theorem 1 is true from $S=B+n H^{2}$.

If $n^{2} H^{2}>4(n-1) c$ and $c \geq 0$ hold, we also have

$$
\begin{gathered}
\sup B-\frac{n-2}{\sqrt{n(n-1)}} n|H| \sup B^{1 / 2}+n c-n H^{2} \\
=\left(\sup B^{1 / 2}-B_{\min }^{1 / 2}\right)\left(\sup B^{1 / 2}-B_{\max }^{1 / 2}\right)
\end{gathered}
$$

(a) When $H^{2} \geq c$, since $B_{\min }^{1 / 2}=\left((n-2) n|H|-n \sqrt{n^{2} H^{2}-4(n-1) c}\right) / 2 \sqrt{n(n-1)} \leq 0$ holds, from (4.9), we infer either sup $B=0$, in this case, $M$ is totally umbilical, or $0<\sup B^{1 / 2} \leq\left((n-2) n|H|+n \sqrt{n^{2} H^{2}-4(n-1) c}\right) / 2 \sqrt{n(n-1)}$.
(b) When $H^{2}<c$, we have $B_{\min }^{1 / 2}=\left((n-2) n|H|-n \sqrt{n^{2} H^{2}-4(n-1) c}\right) / 2 \sqrt{n(n-1)}>$ 0 . Hence, we have, from (4.9), sup $B=0$, in this case, $M$ is totally umbilical, or $B_{\min }^{1 / 2} \leq$ $\sup B^{1 / 2} \leq B_{\text {max }}^{1 / 2}$. Thus, we infer that the assertion (4) in Theorem 1 is true because of $S=B+\bar{n} H^{2}$.
If $S \equiv n / 2(n-1)\left[n^{2} H^{2}-2(n-1) c+(n-2)|H|\left\{n^{2} H^{2}-4(n-1) c\right\}^{1 / 2}\right]$ holds, we know that these inequalities in the proof of Lemma 1 and (4.6) are equalities and $S>n H^{2}$. Hence, we have $n^{2} H^{2} \geq 4(n-1) c$ from (1) in Theorem 1. Thus, we can infer that $n-1$ of the principal curvatures $\lambda_{i}$ are equal. Since the mean curvature $H$ is constant and $S$ is constant, we infer that principal curvatures are constant on $M$. Thus, $M$ is an isoparametric hypersurface with two distinct principal curvatures one of which is simple. This completes the proof of Theorem 1.

Proof of Theorem 2. According to Gauss equation (2.12), we have

$$
R_{j k k j}=R_{j k k j}^{\prime}-\left(h_{j j} h_{k k}-h_{j k} h_{k j}\right)=R_{j k k j}^{\prime}-h_{j j} h_{k k}=R_{j k k j}^{\prime}-\lambda_{j} \lambda_{k}
$$

Hence, we obtain

$$
\begin{equation*}
R_{j k k j}^{\prime}=R_{j k k j}+\lambda_{j} \lambda_{k} \tag{4.10}
\end{equation*}
$$

Therefore

$$
\begin{align*}
& -2 \sum_{i, j, k, l} h_{i j} h_{k l}\left(R_{l i j k}^{\prime}-h_{l i} h_{i j} R_{l k j k}^{\prime}\right) \\
& \quad=-2 \sum_{j, k}\left(\lambda_{j} \lambda_{k} R_{k j j k}^{\prime}-\lambda_{k}^{2} R_{k j k}^{\prime}\right)=-2 \sum_{j, k}\left(\lambda_{j} \lambda_{k}-\lambda_{k}^{2}\right) R_{k j j k}^{\prime}=\sum_{j, k}\left(\lambda_{j}-\lambda_{k}\right)^{2} R_{k j j k}^{\prime} \\
& \quad=\frac{1}{2} \sum_{j, k}\left(\lambda_{j}-\lambda_{k}\right)^{2} R_{k j j k}^{\prime}+\frac{1}{2} \sum_{j, k}\left(\lambda_{j}-\lambda_{k}\right)^{2}\left(R_{k j j k}+\lambda_{j} \lambda_{k}\right) \\
& \quad \geq \frac{c_{2}}{2} \sum_{j, k}\left(\lambda_{j}-\lambda_{k}\right)^{2}+\frac{1}{2} \sum_{j, k}\left(\lambda_{j}-\lambda_{k}\right)^{2}\left(R_{k j j k}+\lambda_{j} \lambda_{k}\right) . \tag{4.11}
\end{align*}
$$

By making use of the same proof as in the proof of Lemma 1, we have

$$
\begin{equation*}
\frac{1}{2} \Delta S \geq\left(n c_{2}+c_{1}\right)\left(S-n H^{2}\right)+\left(S^{2}-n H h_{3}\right)+\frac{1}{2} \sum_{j, k}\left(\lambda_{j}-\lambda_{k}\right)^{2}\left(R_{k j j k}+\lambda_{j} \lambda_{j}\right) \tag{4.12}
\end{equation*}
$$

Since the sectional curvature of $M$ is not less than $-\left(c_{2}+c_{1} / n\right)$, we have

$$
\begin{align*}
& \frac{1}{2} \sum_{j, k}\left(\lambda_{j}-\lambda_{k}\right)^{2}\left(R_{k j j k}+\lambda_{j} \lambda_{i}\right) \\
& \quad \geq-\frac{1}{2}\left(c_{2}+\frac{c_{1}}{n}\right) \sum_{j, k}\left(\lambda_{j}-\lambda_{k}\right)^{2}+\frac{1}{2} \sum_{j, k}\left(\lambda_{j}-\lambda_{k}\right)^{2} \lambda_{j} \lambda_{k} \\
& \quad=-\left(c_{2}+\frac{c_{1}}{n}\right)\left(n S-n^{2} H^{2}\right)+\left(n H h_{3}-S^{2}\right) \tag{4.13}
\end{align*}
$$

Thus, we infer, from (4.12) and (4.13)

$$
\begin{equation*}
\frac{1}{2} \Delta S \geq\left(n c_{2}+c_{1}\right)\left(S-n H^{2}\right)+\left(S^{2}-n H h_{3}\right)+\frac{1}{2} \sum_{j, k}\left(\lambda_{j}-\lambda_{k}\right)^{2}\left(R_{k j j k}+\lambda_{j} \lambda_{i}\right) \geq 0 \tag{4.14}
\end{equation*}
$$

From Theorem 1, we know that $S$ is bounded. Applying the Generalized Maximum Principle to $S$, we have that there exists a point sequence $\left\{p_{m}\right\} \subset M$ such that

$$
\begin{equation*}
\lim \sup _{m \rightarrow \infty} \Delta S\left(p_{m}\right) \leq 0, \quad \lim _{m \rightarrow \infty}|\nabla S|\left(p_{m}\right)=0, \quad \lim _{m \rightarrow \infty} S\left(p_{m}\right)=\sup S \tag{4.15}
\end{equation*}
$$

Hence, we have

$$
\begin{equation*}
\lim _{m \rightarrow \infty} \sum_{j, k}\left(\lambda_{j}-\lambda_{k}\right)^{2}\left\{R_{k j j k}+\left(c_{2}+\frac{c_{1}}{n}\right)\right\}\left(p_{m}\right)=0 \tag{4.16}
\end{equation*}
$$

and

$$
\begin{equation*}
\lim _{m \rightarrow \infty} \sum_{j, k}\left(\lambda_{j}-\lambda_{k}\right)^{2} R_{k j j k}^{\prime}=c_{2} \lim _{m \rightarrow \infty} \sum_{j, k}\left(\lambda_{j}-\lambda_{k}\right)^{2} . \tag{4.17}
\end{equation*}
$$

Since the function $S=\sum_{j} \lambda_{j}^{2}$ is bounded, $\left\{\lambda_{j}\left(p_{m}\right)\right\}$ is a bounded sequence for any $j$. Thus, we can assume

$$
\begin{equation*}
\lim _{m \rightarrow \infty} \lambda_{j}\left(p_{m}\right)=\lambda_{j_{0}} \tag{4.18}
\end{equation*}
$$

for any $j$, if necessary, we can take a subsequence of $\left\{\lambda_{j}\left(p_{m}\right)\right\}$. Hence, from (4.10) and (4.17), we have

$$
\begin{equation*}
\lim _{m \rightarrow \infty} \sum_{j, k}\left(\lambda_{j}-\lambda_{k}\right)^{2} R_{k j j k}\left(p_{m}\right)=\lim _{m \rightarrow \infty} \sum_{j, k}\left(\lambda_{j}-\lambda_{k}\right)^{2}\left(c_{2}-\lambda_{j} \lambda_{k}\right)\left(p_{m}\right) . \tag{4.19}
\end{equation*}
$$

Therefore, we infer, from (4.16) and (4.19)

$$
\begin{equation*}
\lim _{m \rightarrow \infty} \sum_{j, k}\left(\lambda_{j}-\lambda_{k}\right)^{2}\left(c-\lambda_{j} \lambda_{i}\right)=0 \tag{4.20}
\end{equation*}
$$

If $\lambda_{j_{0}} \neq \lambda_{k_{0}}$, from (4.17), we have

$$
\begin{aligned}
\lim _{m \rightarrow \infty}\left(c_{2}-\lambda_{j} \lambda_{k}\right)\left(p_{m}\right) & =\lim _{m \rightarrow \infty}\left(R_{k j j k}^{\prime}-\lambda_{j} \lambda_{k}\right)\left(p_{m}\right) \\
& =\lim _{m \rightarrow \infty} R_{k j j k}\left(p_{m}\right) \geq-\left(c_{2}+\frac{c_{1}}{n}\right) .
\end{aligned}
$$

Hence, in this case, we obtain

$$
\begin{equation*}
\lim _{m \rightarrow \infty}\left(c-\lambda_{j} \lambda_{i}\right)\left(p_{m}\right) \geq 0 . \tag{4.21}
\end{equation*}
$$

Thus, (4.20) and (4.21) yield, for any $i$ and $j$

$$
\begin{equation*}
\lim _{m \rightarrow \infty}\left(-\lambda_{i} \lambda_{j}\left(p_{m}\right)+c\right)\left(\lambda_{i}-\lambda_{j}\right)^{2}\left(p_{m}\right)=0 \tag{4.22}
\end{equation*}
$$

By (4.18) and (4.22) we get

$$
\begin{equation*}
\left(-\lambda_{i_{0}} \lambda_{j_{0}}+c\right)\left(\lambda_{i_{0}}-\lambda_{j_{0}}\right)^{2}=0 \tag{4.23}
\end{equation*}
$$

for any $i_{0}$ and $j_{0}$. By the simple algebraic calculation, it is clear that at most two of $\left\{\lambda_{j_{0}}\right\}$ 's are distinct. If all of $\left\{\lambda_{j_{0}}\right\}$ 's coincides with each other, then, from (4.21), we have $c \geq \lambda_{j_{0}}^{2} \geq 0$.

If two of the $\left\{\lambda_{j_{0}}\right\}$ 's are distinct, which are denoted by $\lambda$ and $\mu(\lambda \neq \mu)$. By (4.23), they satisfy

$$
\begin{equation*}
-\lambda \mu+c=0 \tag{4.24}
\end{equation*}
$$

Now let us denote by $r$ and $s$ the number of indices $\lambda_{j}\left(p_{m}\right) \rightarrow \lambda$ and $\lambda_{j}\left(p_{m}\right) \rightarrow \mu$, respectively. Then we want to assert that $r=1$ or $s=1$.

In fact, if $r \geq 2$ and $s \geq 2$ hold, it follows from $r \geq 2$ that there are distinct indices $i$ and $j$ such that $\lambda_{i}\left(p_{m}\right) \rightarrow \lambda$ and $\lambda_{j}\left(p_{m}\right) \rightarrow \lambda(m \rightarrow \infty)$ and hence we have

$$
\lim _{m \rightarrow \infty}\left(-\lambda_{i} \lambda_{j}+c\right)=-\lambda^{2}+c, \quad \text { for } i \neq j
$$

By (4.18) and (4.21), we obtain $c \geq \lambda^{2} \geq 0$. Similarly, we have $c \geq \mu^{2} \geq 0$, which implies that

$$
c^{2} \geq \lambda^{2} \mu^{2}
$$

On the other hand, since we see that $c=\lambda \mu$ in (4.24), it implies $\lambda^{2}=\mu^{2}=c$. Furthermore, it turns out to be $\lambda= \pm \mu$. Because they are distinct, it yields $c=\lambda \mu=-\lambda^{2}$. Hence we obtain $\lambda=\mu=0$. This is impossible because of $\lambda \neq \mu$. Thus, our assertion is true.

Without loss of generality, we assume $r=1$. Since $n>2$, we have $s \geq 2$ and by the above discussion we obtain $0<\mu^{2}<c$. This finishes the proof of the first part of assertions in Theorem 2.

Next, we shall prove the second part of the assertions in Theorem 2. From the above assertion, together with (4.24) it follows that $\lambda^{2}>c$. It is clear that we may assume that the mean curvature $H$ is positive. Then $\lambda$ and $\mu$ are positive because of $c>0$ and (4.24). By defining $c_{3}$ and $c_{4}$ by $\lambda^{2}=c-c_{3}$ and $\mu^{2}=c-c_{4}$, respectively, we have $c_{3}<0,0<c_{4}<c$ and

$$
\begin{equation*}
\left(c-c_{3}\right)\left(c-c_{4}\right)=c^{2}, \quad \text { i.e., } \quad \frac{1}{c_{3}}+\frac{1}{c_{4}}=\frac{1}{c} \tag{4.25}
\end{equation*}
$$

Since the mean curvature $H$ is constant, we obtain

$$
n H=\lambda+(n-1) \mu=\left(c-c_{3}\right)^{1 / 2}+(n-1)\left(c-c_{4}\right)^{1 / 2}, \quad \lambda \mu=c
$$

Hence, we have

$$
\lambda^{2}-n H \lambda+(n-1) c=0
$$

Since $H^{2} \geq c$ holds, we conclude

$$
\begin{equation*}
\lambda=\frac{c}{\mu}=\frac{n H+\sqrt{n^{2} H^{2}-4(n-1) c}}{2}, \quad \text { and } \quad \mu=\frac{n H-\sqrt{n^{2} H^{2}-4(n-1) c}}{2(n-1)} . \tag{4.26}
\end{equation*}
$$

Hence, we infer

$$
\begin{align*}
S & =\lambda^{2}+(n-1) \mu^{2} \\
& =\frac{1}{2(n-1)}\left[n\left\{n^{2} H^{2}-2(n-1) c\right\}+(n-2) n|H|\left\{n^{2} H^{2}-4(n-1) c\right\}^{1 / 2}\right] \tag{4.27}
\end{align*}
$$

It is a contradiction. Therefore

$$
\lambda_{j_{0}}=\lambda
$$

for any $j$. This implies

$$
\sup S=n H^{2}
$$

From $S-n H^{2}=B \geq 0$, we have $S \equiv n H^{2}$. Hence, $M$ is totally umbilical.

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## References

[1] K. Akutagawa, On space-like hypersurfaces with constant mean curvature in the de Sitter space, Math. Z. 196 (1987) 12-19.
[2] E. Calabi, Examples of Bernstein problems for some nonlinear equations, Proc. Pure Appl. Math. 15 (1970) 223-230.
[3] Q.-M. Cheng, Complete space-like submanifolds in a de Sitter space with parallel mean curvature vector, Math. Z. 206 (1991) 333-339.
[4] Q.-M. Cheng, Hypersurfaces in a Lorentz space form, Arch. Math. 63 (1994) 271-281.
[5] Q.-M. Cheng, S. Ishikawa, Space-like hypersurfaces with constant scalar curvature, Manus. Math. 95 (1998) 499-505.
[6] Q.-M. Cheng, H. Nakagawa, Totally umbilic hypersurfaces, Hiroshima Math. J. 20 (1990) 1-10.
[7] S.Y. Cheng, S.T. Yau, Maximal space-like hypersurfaces in the Lorentz-Minkovski spaces, Ann. Math. 104 (1976) 223-230.
[8] S.M. Choi, J.-H. Kwon, Y.J. Suh, A Liouville type theorem for complete Riemannian manifolds, Bull. Korean Math. Soc. 35 (1998) 301-309.
[9] S.M. Choi, S.M. Lyu, Y.J. Suh, Complete space-like hypersurfaces in a Lorentz manifold, Math. J. Toyama Univ. 22 (1999) 53-76.
[10] Y. Chouque-Bruhat, A.E. Fisher, J.E. Marsdan, Maximal hypersurfaces and positivity mass, in: J. Ehlers (Ed.), Proceedings of the E. Fermi Summer School of the Italian Physical Society, North-Holland, Amsterdam, 1979.
[11] T. Ishihara, Maximal space-like submanifolds of a pseudo-Riemannian space of constant curvature, Mich. Math. J. 35 (1988) 345-352.
[12] U.-H. Ki, H.-J. Kim, H. Nakagawa, On space-like hypersurfaces with constant mean curvature of a Lorentz space form, Tokyo J. Math. 14 (1991) 205-216.
[13] J. Marsdan, F. Tipler, Maximal hypersurfaces and foliations of constant mean curvature in general relativity, Bull. Am. Phys. Soc. 23 (1978) 84.
[14] S. Montiel, An integral inequality for compact space-like hypersurfaces in de Sitter space and applications to the case of constant mean curvature, Indiana Univ. Math. J. 37 (1988) 909-917.
[15] S. Nishikawa, On maximal spacelike hypersurfaces in a Lorenzian manifolds, Nagoya Math. J. 95 (1984) 117-124.
[16] H. Omori, Isometric immersions of Riemannian manifolds, J. Math. Soc. Jpn. 19 (1967) 205-211.
[17] B. O'Neill, Semi-Riemannian Geometry with Applications to Relativity, Academic Press, New York, 1983.
[18] J. Ramanathan, Complete space-like hypersurfaces of constant mean curvature in de Sitter space, Indiana Univ. Math. 36 (1987) 349-359.
[19] S. Stumbles, Hypersurfaces of constant mean extrinsic curvature, Ann. Phys. 133 (1980) 28-56.
[20] Y.J. Suh, Y.S. Choi, H.Y. Yang, On space-like hypersurfaces with constant mean curvature in a Lorentz manifold, Houston J. Math. 28 (2002) 47-70.
[21] S.T. Yau, Harmonic functions on complete Riemannian manifolds, Commun. Pure Appl. Math. 28 (1975) 201-228.


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